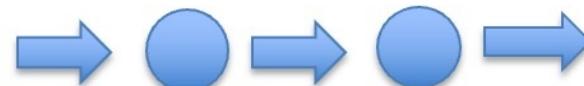
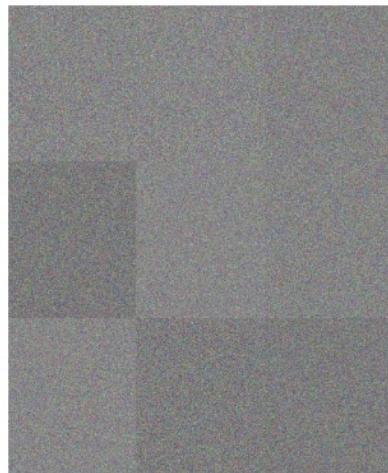




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HOLLYWOOD STARS



APS DFD23 WASHINGTON Synthetic Lagrangian Turbulence by Generative Diffusion Models

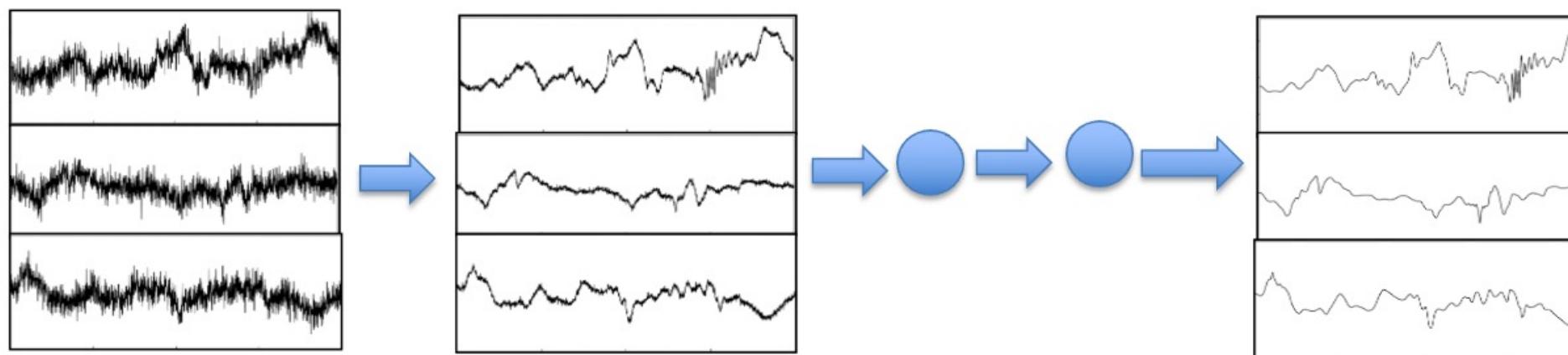
T. Li, LB, F. Bonaccorso, M. Scarpolini, M. Buzzicotti.
Synthetic Lagrangian Turbulence by Generative Diffusion Models.
[arXiv:2307.08529 \(2023\)](https://arxiv.org/abs/2307.08529) – Submitted to Nature Machine Intelligence

CREDITS: T. LI, F. BONACCORSO, M. BUZZICOTTI, M. SCARPOLINI



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APS DFD23 WASHINGTON Synthetic Lagrangian Turbulence by Generative Diffusion Models

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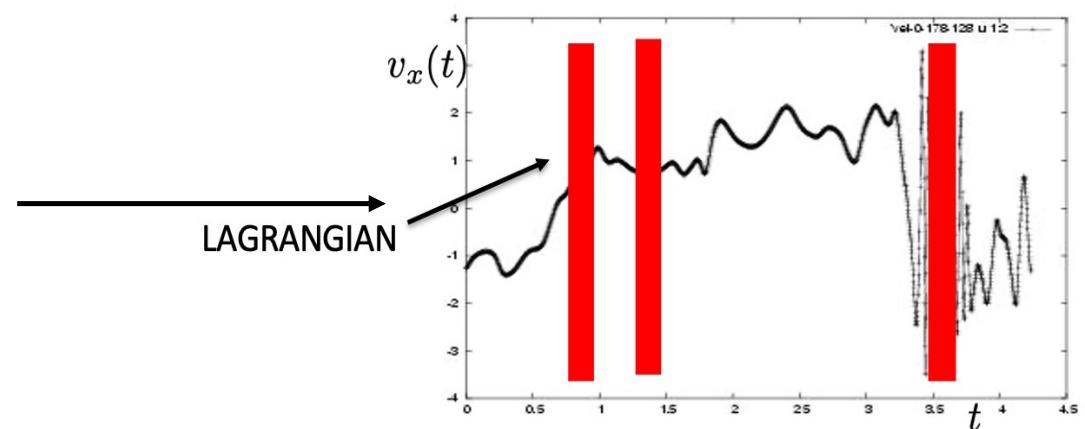
STOCHASTIC MODELS FOR LAGRANGIAN TURBULENCE: WHY?

T. Li, LB, F. Bonaccorso, M. Scarpolini, M. Buzzicotti.
Synthetic Lagrangian Turbulence by Generative Diffusion Models.
arXiv:2307.08529 (2023) – Submitted to Nature Machine Intelligence

GENERATION OF LARGE SYNTHETIC DATA-BASE FOR

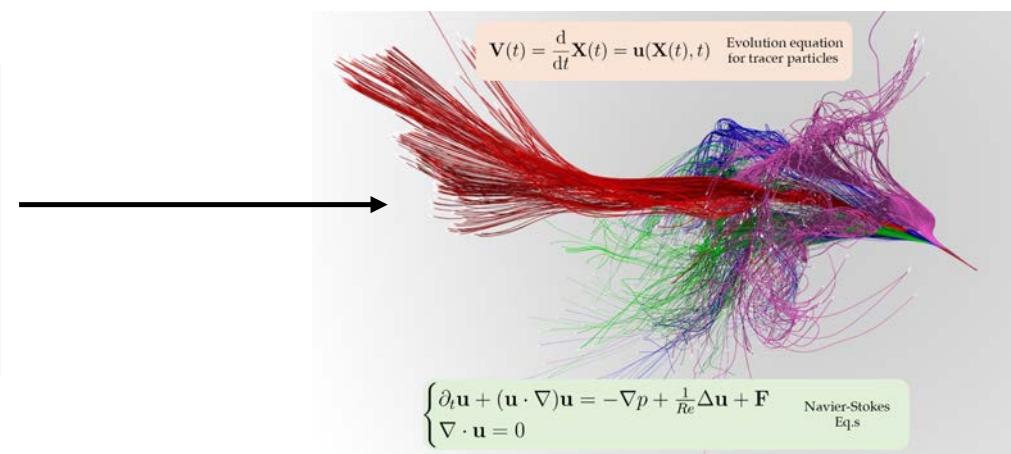
- (I) RANKING OF PHYSICS FEATURES
- (II) TESTING DOWNSTREAM APPLICATIONS/MODELS

DATA ASSIMILATION/INPAINTING FROM MISSING FIELD/EXPERIMENTAL OBSERVATION

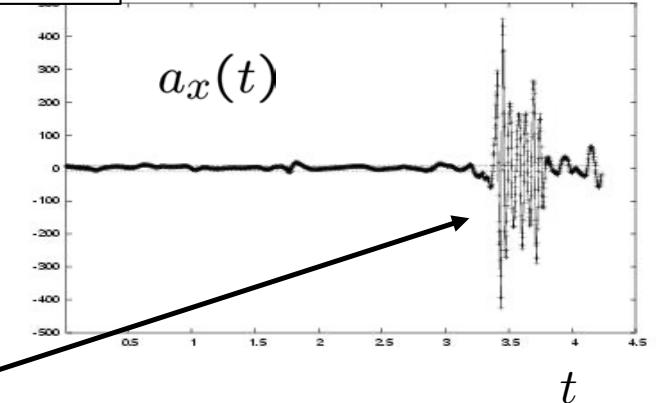
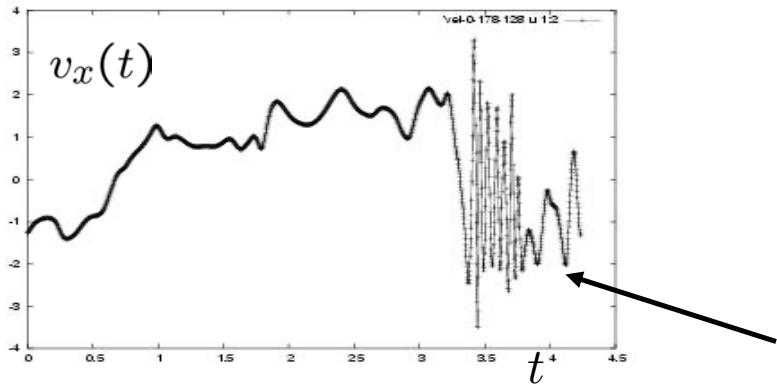


CLASSIFICATION/INFERRAL OF MISSING/INTERNAL PROPERTIES:

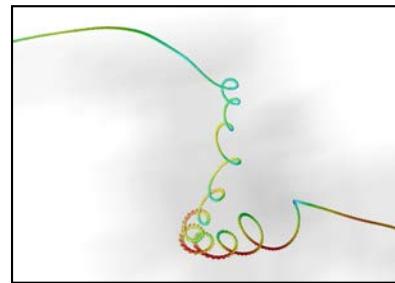
- (I) INERTIA
- (II) SHAPE
- (III) ACTIVE DEGREES OF FREEDOM
- (IV)



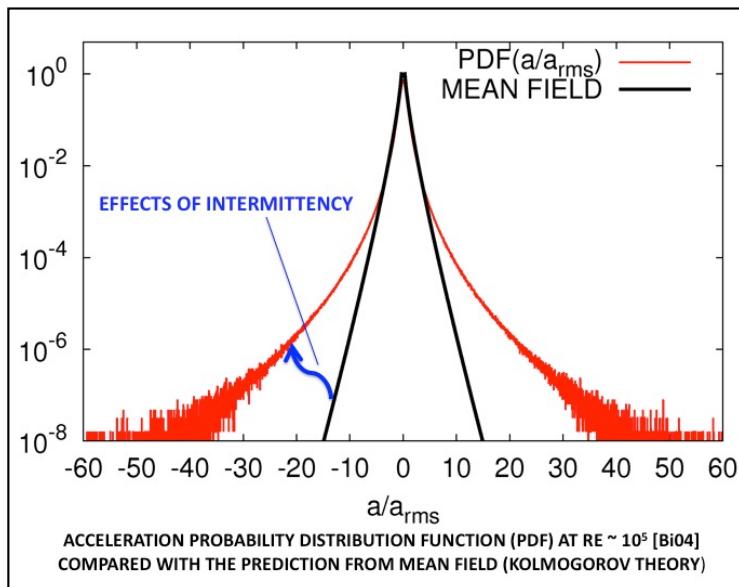
$$\begin{cases} \mathbf{a} = \partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = -\nabla p + \nu \Delta \mathbf{u} + \mathbf{f} \\ \nabla \cdot \mathbf{u} = 0 \end{cases}$$



EXTREME EVENTS



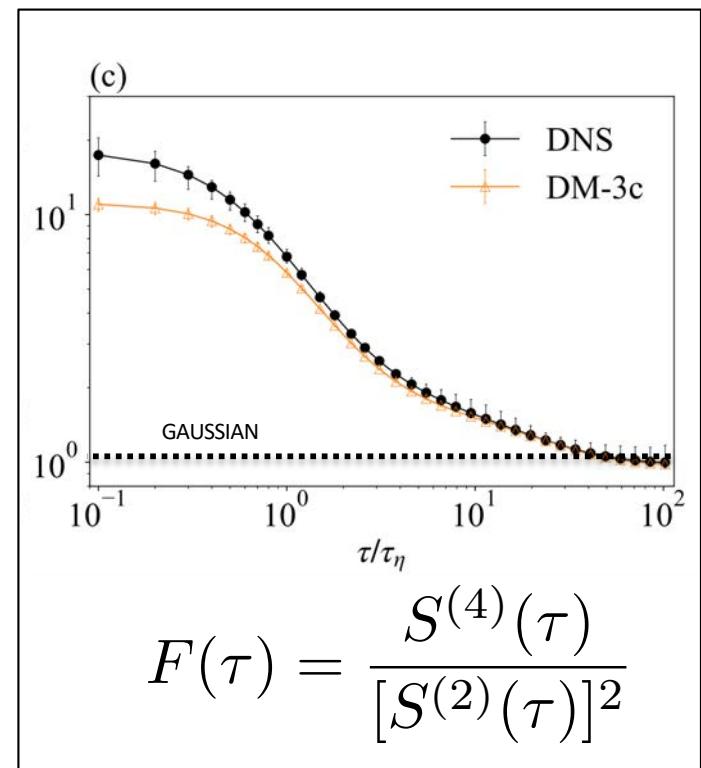
$$S_i^{(p)}(\tau) = \langle [v_i(t + \tau) - v_i(t)]^p \rangle$$



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$$F(\tau) = \frac{S^{(4)}(\tau)}{[S^{(2)}(\tau)]^2}$$

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Diffusion Models

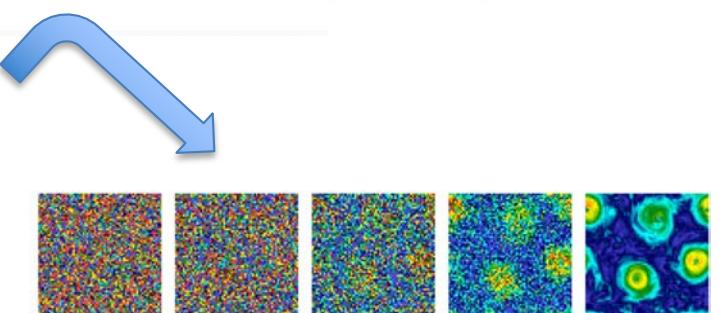
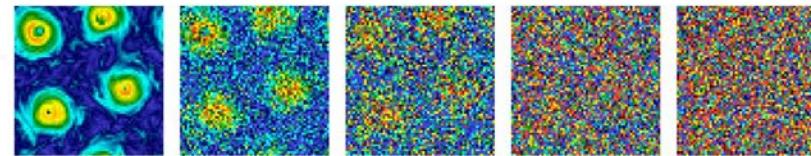
Training set: a set of images $\vec{a}^\mu \in \mathbb{R}^N$ $\mu = 1, \dots, P$
N is the dimension of the data, P their number

Langevin equation for an Ornstein-Uhlenbeck process

$$\frac{d\vec{x}}{dt} = -\vec{x} + \vec{\eta}(t) \quad \langle \eta_i(t)\eta_j(t') \rangle = 2T\delta_{ij}\delta(t-t')$$

$\vec{x}^\mu(t=0) = \vec{a}^\mu$ It transforms the data in iid Gaussian $\mathcal{N}(0, 1)$ at $t \gg 1$

$$P_t(\vec{x}) = \int d\vec{a} P_0(\vec{a}) \frac{1}{\sqrt{2\pi\Delta_t^N}} \exp\left(-\frac{1}{2} \frac{(\vec{x} - \vec{a}e^{-t})^2}{\Delta_t}\right) = \int d\vec{a} P_t(\vec{a}, \vec{x})$$
$$\Delta_t = T(1 - e^{-2t})$$



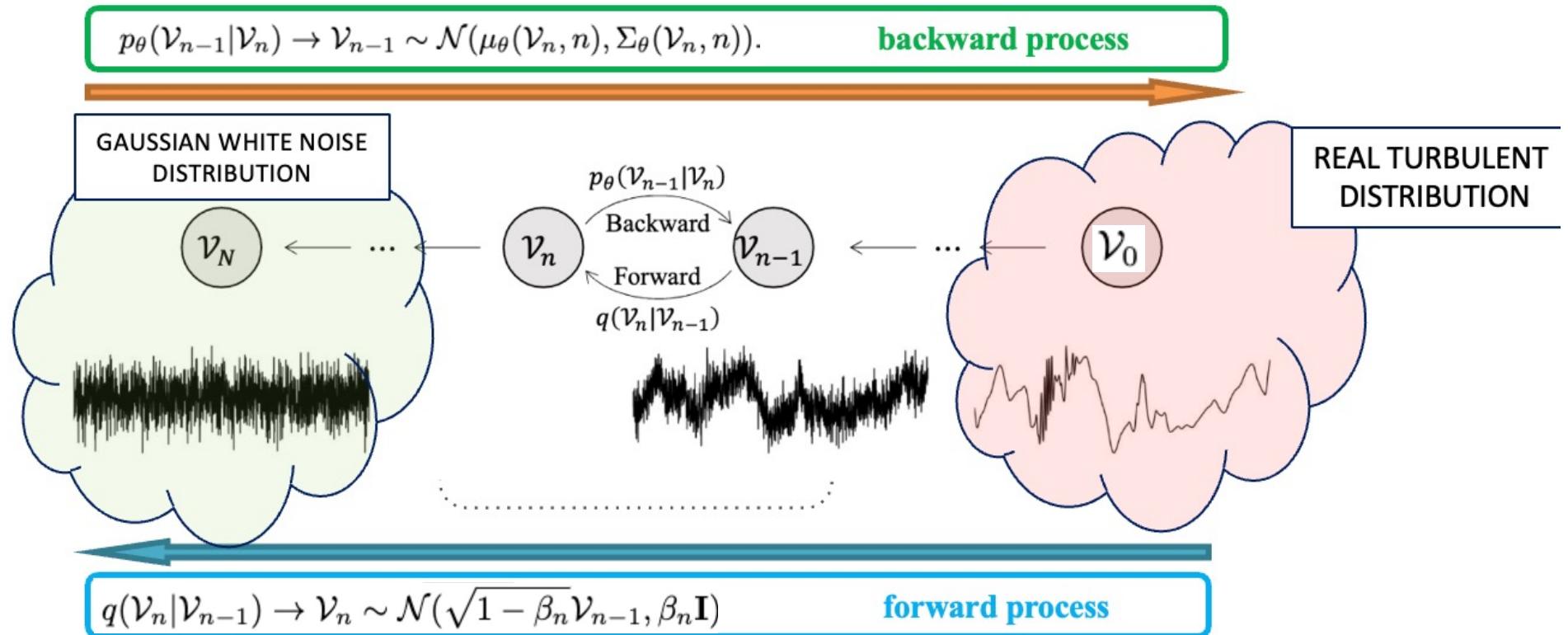
Score function provides the force field to go back in time

$$\mathcal{F}_i(\vec{x}, t) = \frac{\partial \log P_t(\vec{x})}{\partial x_i}$$

$$-\frac{dy_i}{dt} = y_i + 2T\mathcal{F}_i(y, t) + \eta_i(t)$$

Diffusion Models

‘Synthetica Lagrangian Turbulence: all you need is Diffusion Models’ T. Li, L.B, F. Bonaccorso, M. Scarpolini and M. Buzzicotti (arXiv:2307.08529 2023, submitted Nature Machine Intelligence)

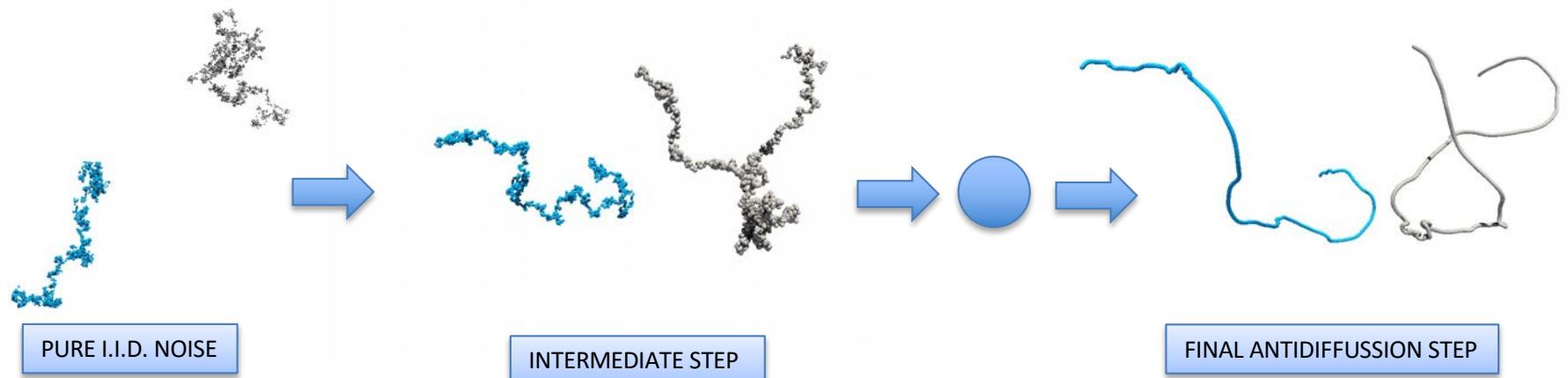
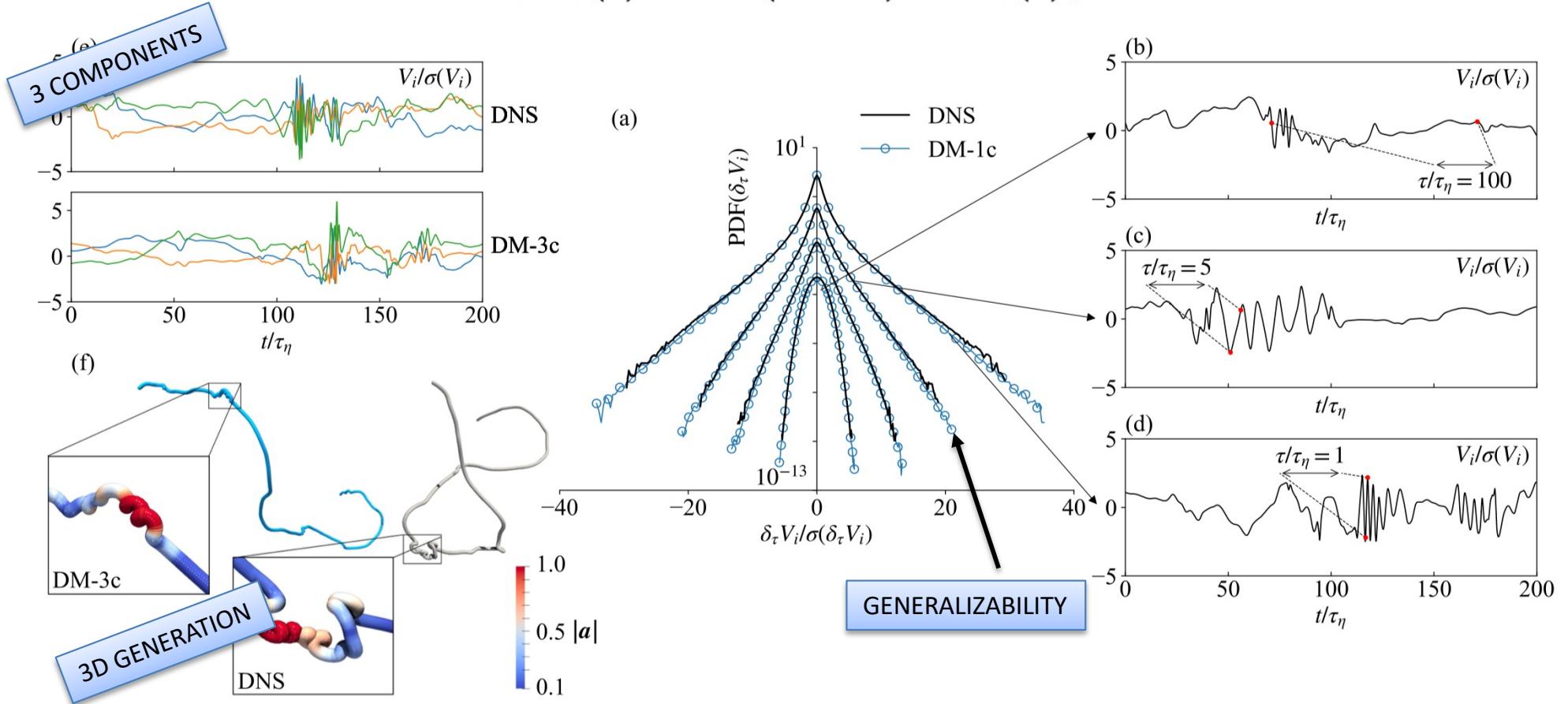


[Sohl-Dickstein et al., Deep Unsupervised Learning using Nonequilibrium Thermodynamics, ICML 2015](#)

[Ho et al., Denoising Diffusion Probabilistic Models, NeurIPS 2020](#)

[Song et al., Score-Based Generative Modeling through Stochastic Differential Equations, ICLR 2021](#)

$$\delta_\tau V_i(t) = V_i(t + \tau) - V_i(t),$$

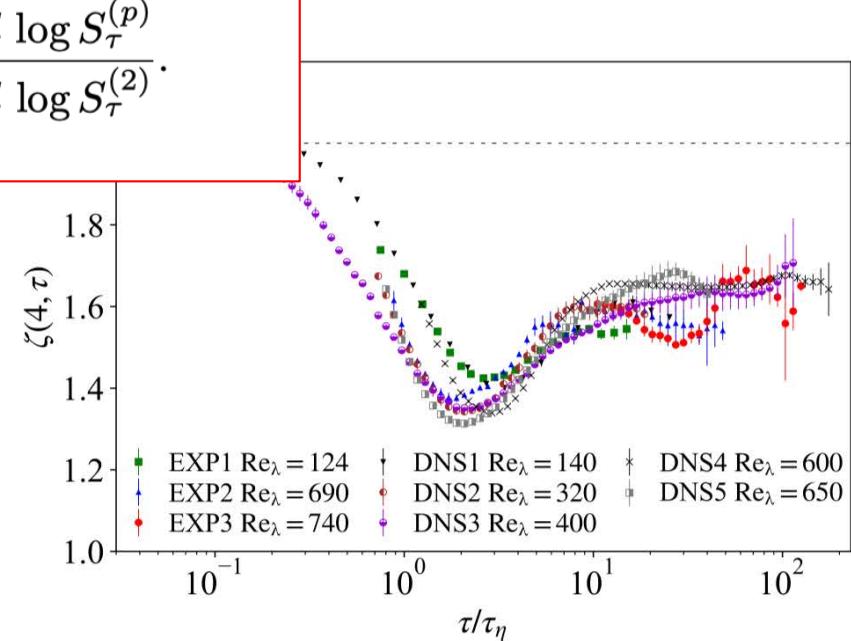
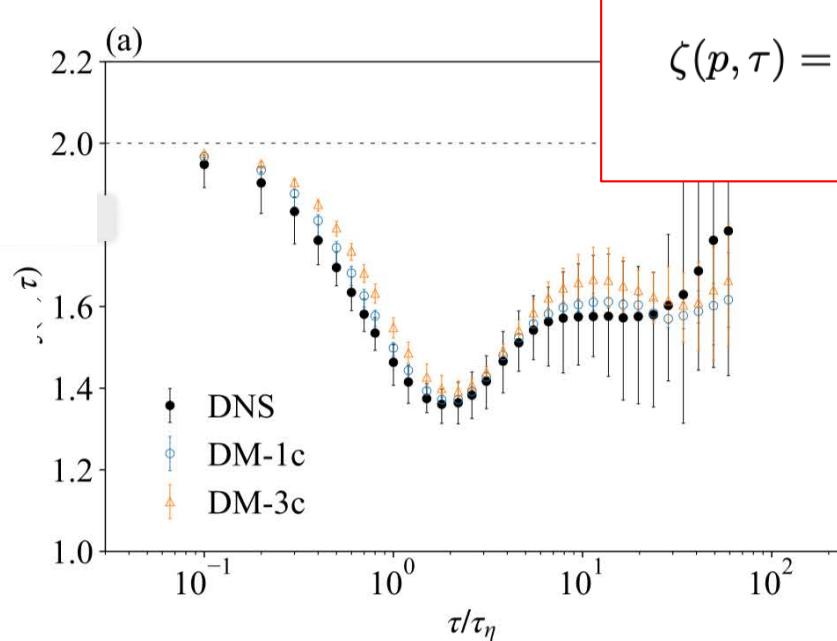
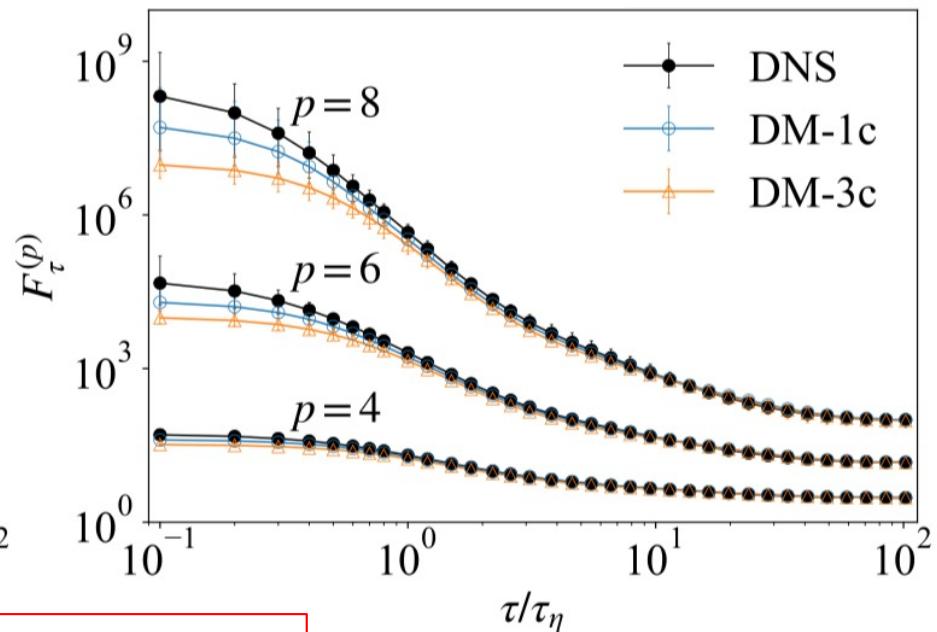
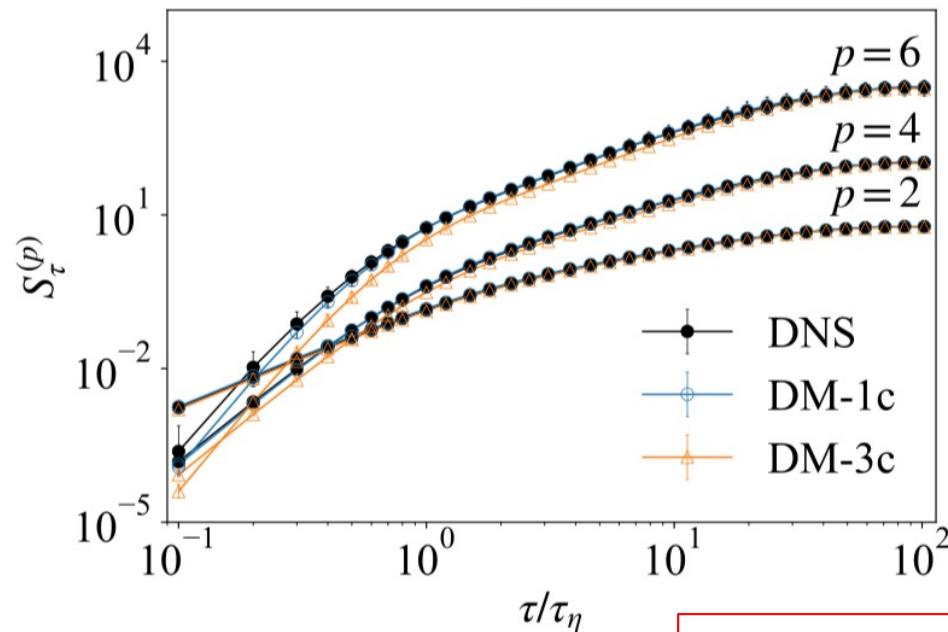


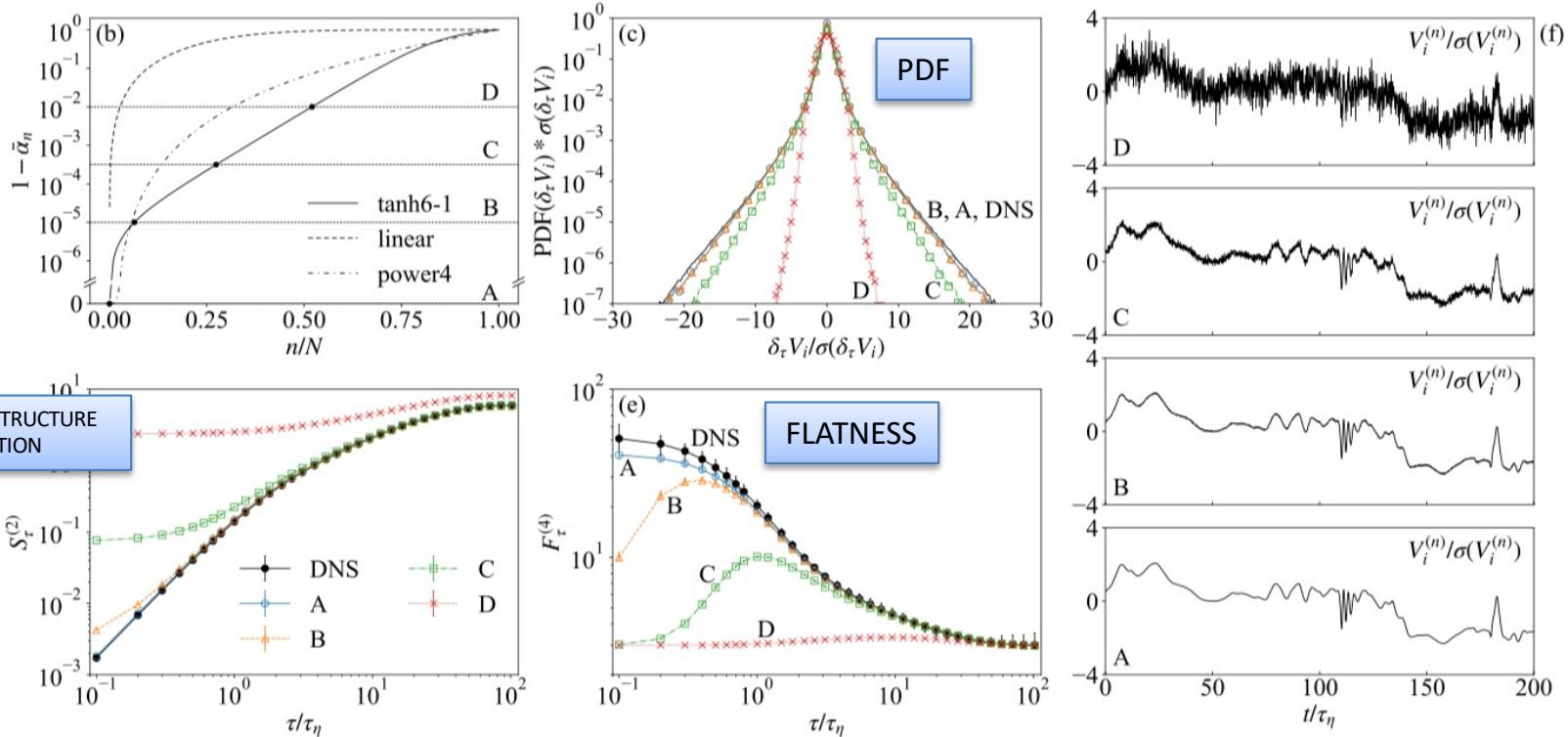
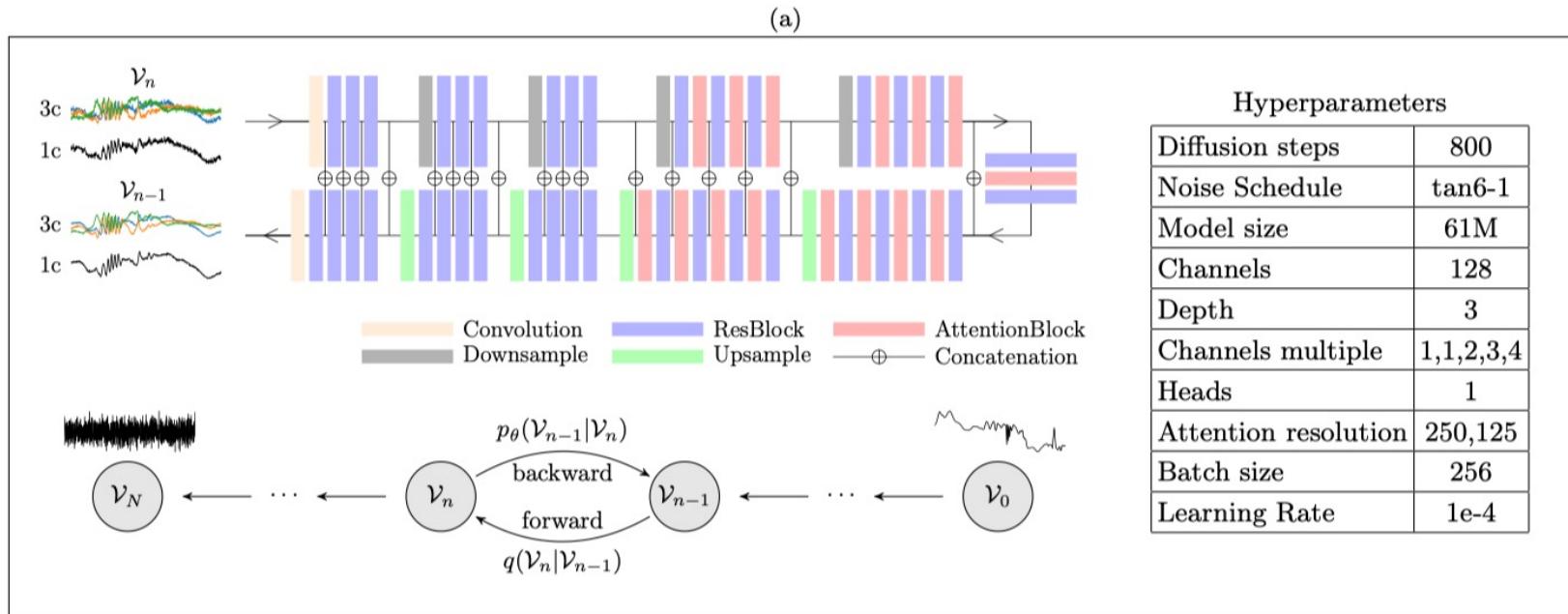
LAGRANGIAN STRUCTURE FUNCTIONS

GENERALIZED FLATNESS

$$S_\tau^{(p)} = \langle (\delta_\tau V_i)^p \rangle$$

$$F_\tau^{(p)} = S_\tau^{(p)} / [S_\tau^{(2)}]^{p/2}.$$





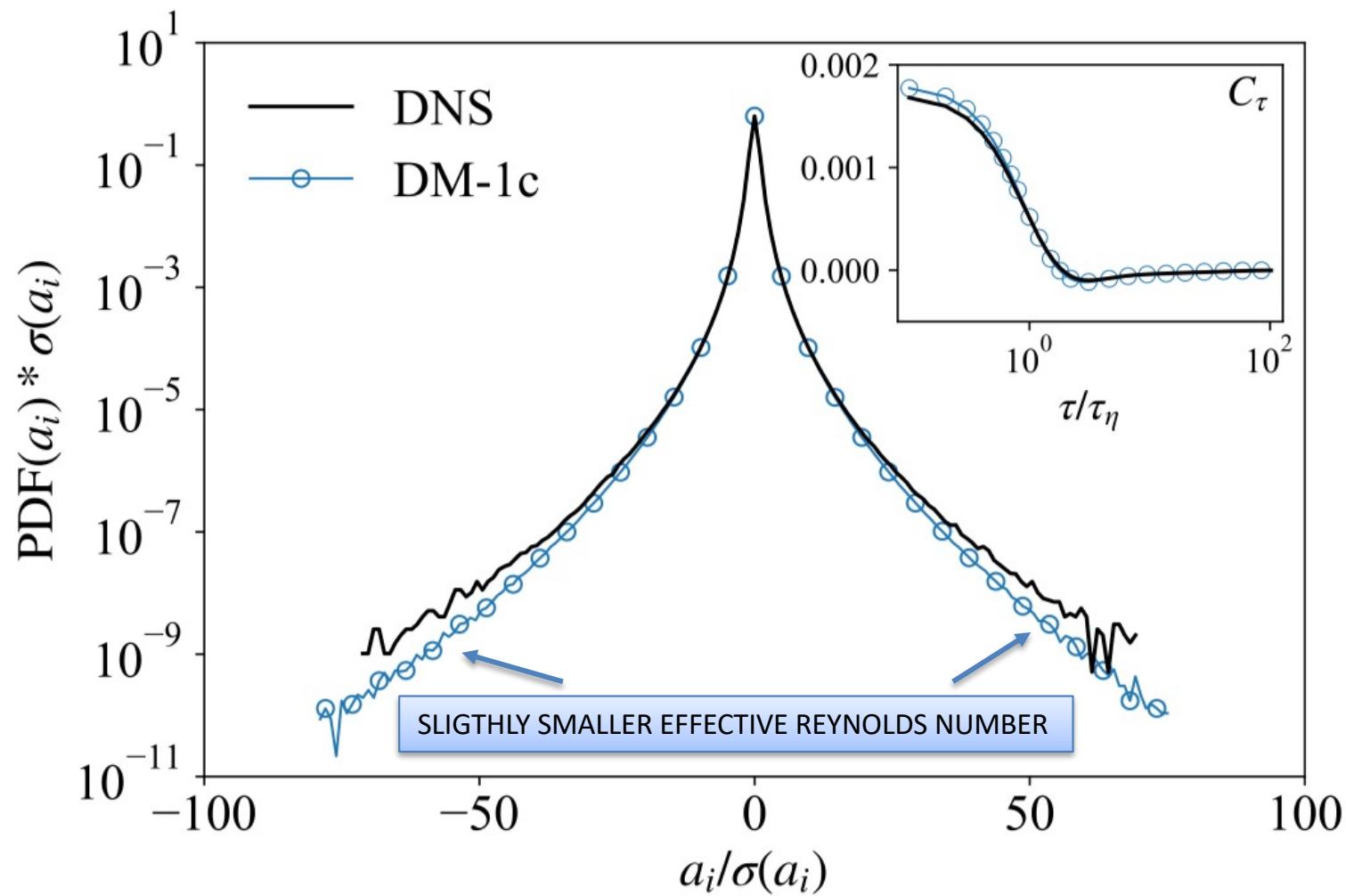
WHAT WE HAVE:

- QUICK STOCHASTIC TOOL TO GENERATE REALISTIC 3D TRAJECTORIES OF TRACERS IN HOMOGENEOUS AND ISOTROPIC TURBULENCE, EASY TO GENERALISE FOR DIFFERENT APPLICATIONS
- IMPRESSIVE QUANTITATIVE AGREEMENT WITH MULTI-SCALE STATISTICAL PROPERTIES

WHAT WE MISS:

- UNDERSTADING OF ROBUSTNESS IN GENERALISING OUT-OF-SAMPLE:
EXTREME EVENTS, DIFFERENT REYNOLDS NUMBERS, DIFFERENT PARTICLES' PROPERTIES
- UNDERSTANDING SCALING PROPERTIES FOR TIME-TO-SOLUTION AT CHANGING IN-SAMPLE PROPERTIES, I.E. AT CHANGING DIMENSION OF THE TRAINING DATASET,
SETS OF HYPER-PARAMETERS, CNN ARCHITECTURES: GAN, DM, TRANSFORMERS
- WHAT-IF QUESTIONS: EXPLICABILITY OF THE GENERATED DATA, FEATURES RANKINGS,
PHYSICS DISCOVERY

ACCELERATION PDF



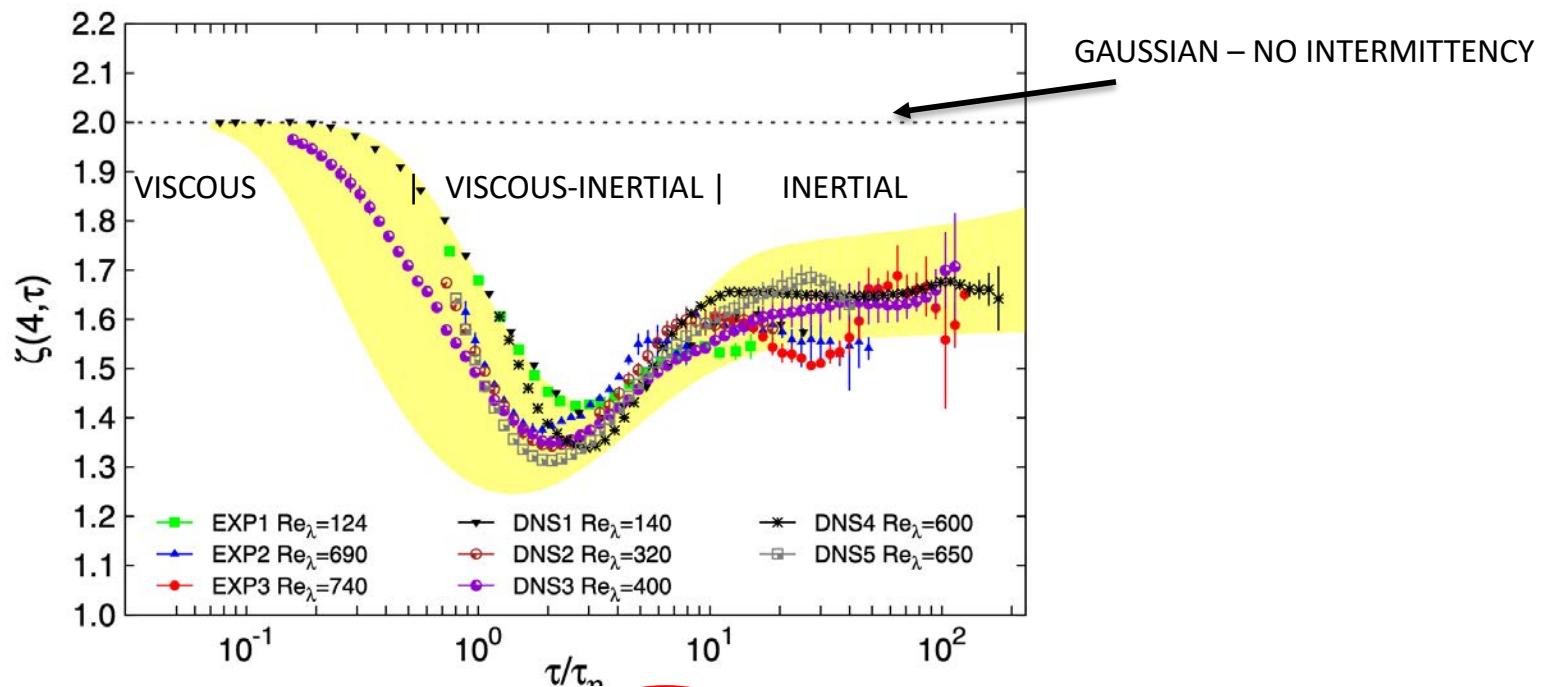
$$S_i^{(p)}(\tau) = \langle [v_i(t + \tau) - v_i(t)]^p \rangle$$

$$\zeta(4, \tau) = \frac{d \log S^{(4)}(\tau)}{d \log S^{(2)}(\tau)}$$

Universal Intermittent Properties of Particle Trajectories in Highly Turbulent Flows

A. Arnéodo,¹ R. Benzi,² J. Berg,³ L. Biferale,^{4,*} E. Bodenschatz,⁵ A. Busse,⁶ E. Calzavarini,⁷ B. Castaing,¹ M. Cencini,^{8,*} L. Chevillard,¹ R. T. Fisher,⁹ R. Grauer,¹⁰ H. Homann,¹⁰ D. Lamb,⁹ A. S. Lanotte,^{11,*} E. Lévèque,¹ B. Lüthi,¹² J. Mann,³ N. Mordant,¹³ W.-C. Müller,⁶ S. Ott,³ N. T. Ouellette,¹⁴ J.-F. Pinton,¹ S. B. Pope,¹⁵ S. G. Roux,¹ F. Toschi,^{16,17,*} H. Xu,⁵ and P. K. Yeung¹⁸

(International Collaboration for Turbulence Research)



$$\delta_\tau v(h) = V_0 \frac{\tau}{T_L} \left[\left(\frac{\tau}{T_L} \right)^\beta + \left(\frac{\tau_\eta}{T_L} \right)^\beta \right]^{(2h-1)/\beta(1-h)}, \quad \tau_\eta(h)/T_L \sim R_\lambda^{2(h-1)/(1+h)}$$

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Guide for users

TURB-Rot. A LARGE DATABASE OF 3D AND 2D SNAPSHOTTS FROM TURBULENT ROTATING FLOWS

A PREPRINT

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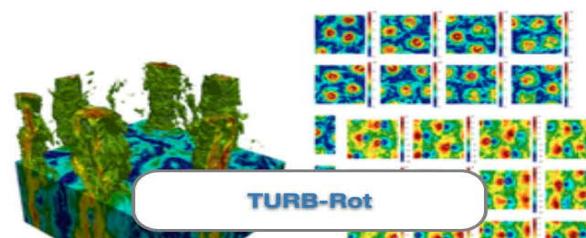


1

Datasets

TURB-Rot

A large database of 3d and 2d snapshots from turbulent rotating



2

Organizations

web_admin

web_admin group

1

member

Organizations	
web_admin	1 member

<https://smart-turb.roma2.infn.it/>

DEAR CHALRES, THE INFLUENCE OF YOUR WORK ON
MY SCIENTIFIC CAREER WENT WELL BEYOND THE
FEW INTERMITTENT SCIENTIFIC COLLABORATION WE
HAD TOGETHER.

MY SCIENTIFIC TRAJECTORY WOULD HAVE NOT BEEN
THE SAME WITHOUT YOUR SCIENCE, YOUR WORK
FOR THE COMMUNITY, YOUR SCIENTIFIC VISION AND
YOUR CONTINUOS SUPPORT.

I AM PROUD CONSIDER MYSELF A GOOD FRIEND OF
YOU AND OF YOUR WONDERFUL FAMILY, HAPPY
BIRTHDAY AND AD MAIORA!

PS: I AM LOOKING FORWARD FOR THE NEXT
BIRTHDAYS,



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Annual Review of Fluid Mechanics